



# **MICROWAVE SPECTROSCOPY AND THE REMOTE SENSING PROBLEM**

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# OVERVIEW

## The Rotation Vibration Spectra of Molecules

Microwave and Infrared Data

Weighted Fits and the Spectral Data Base

## Linewidths, Lineshapes, and the Continua

Physics and Parameterization

Experimental Issues

# SPECTRAL REGIONS

## HISTORICALLY:

MICROWAVE: Low J, Ground vibrational state

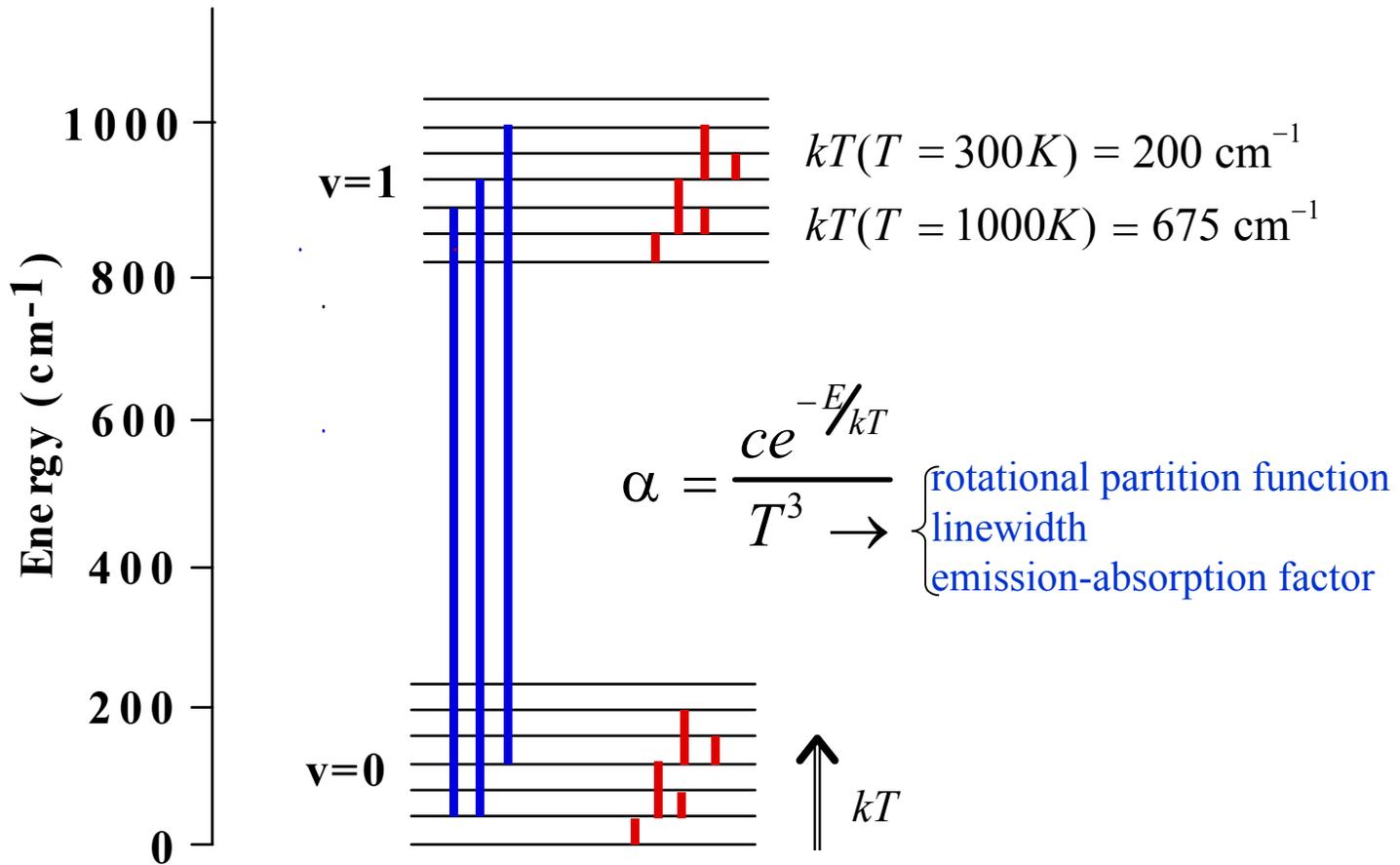
INFRARED: Vibrational bands

## CURRENTLY:

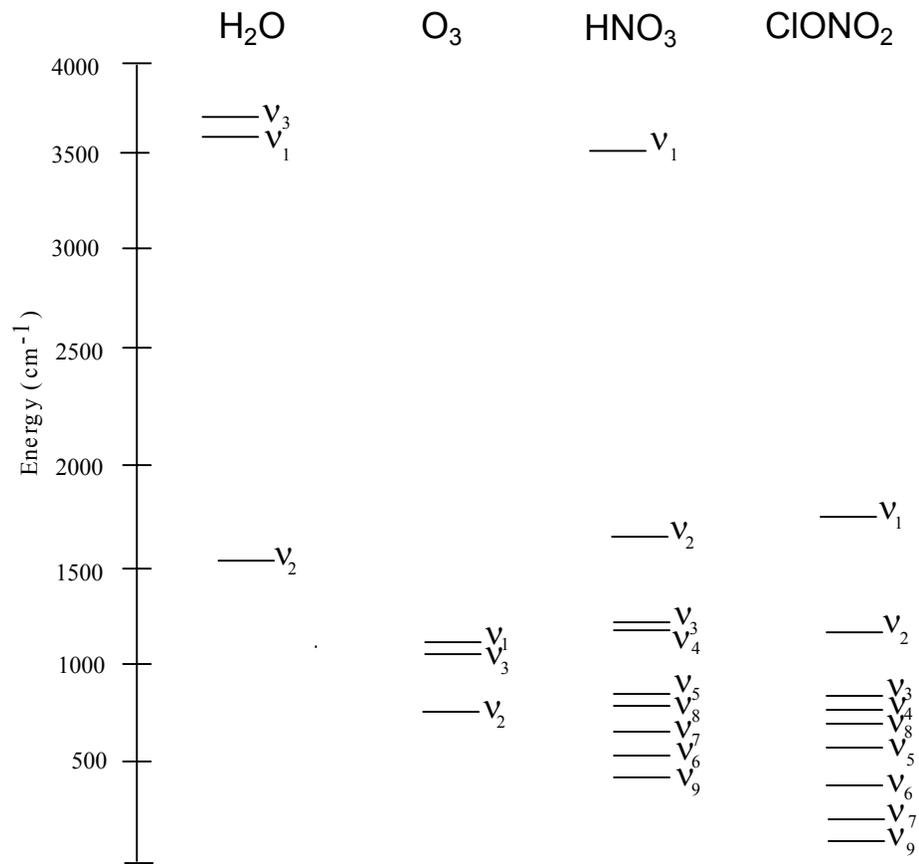
MICROWAVE: Thermally populated rotational manifold in many vibrational states

INFRARED: Doppler limited rotation-vibration spectra

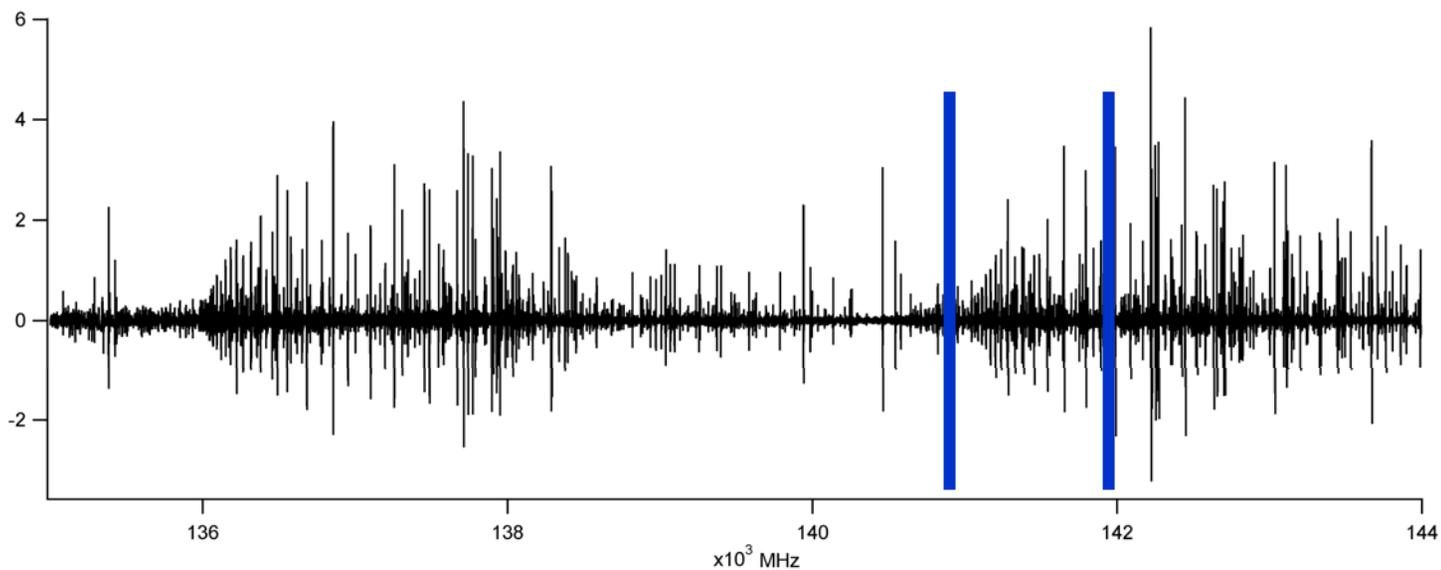
# THE ROTATION-VIBRATION SPECTRUM



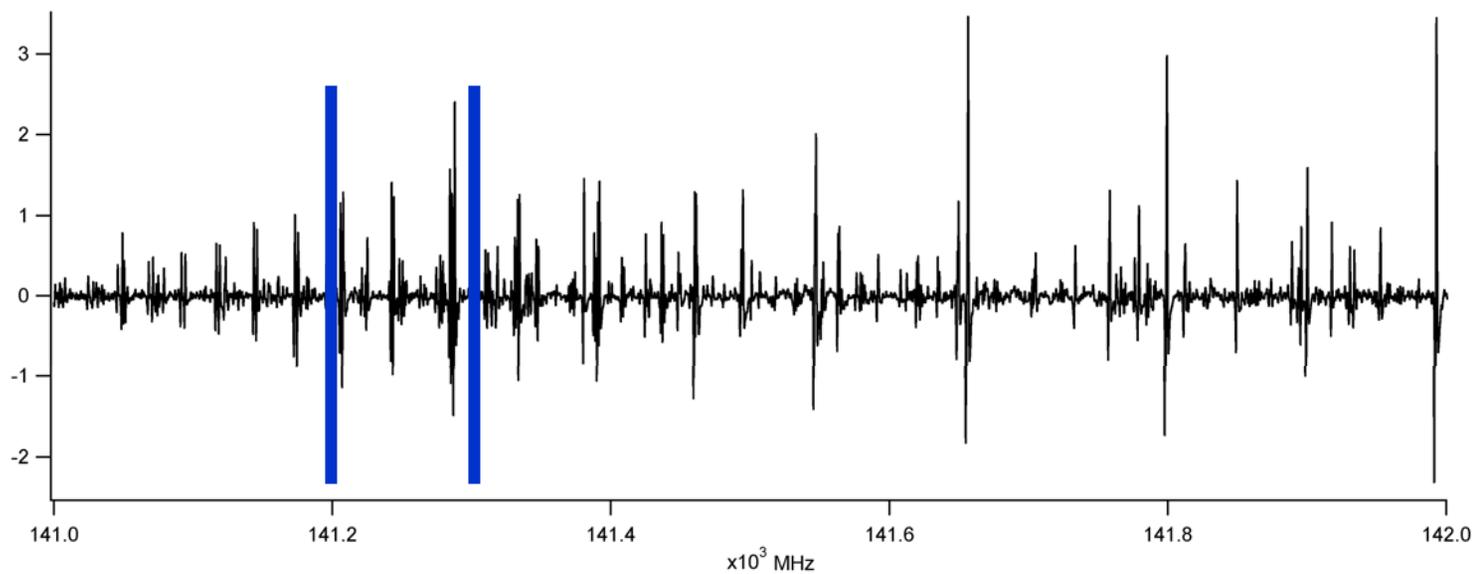
# SPECTRAL DENSITY



# FASSST SPECTRUM OF ClONO<sub>2</sub>

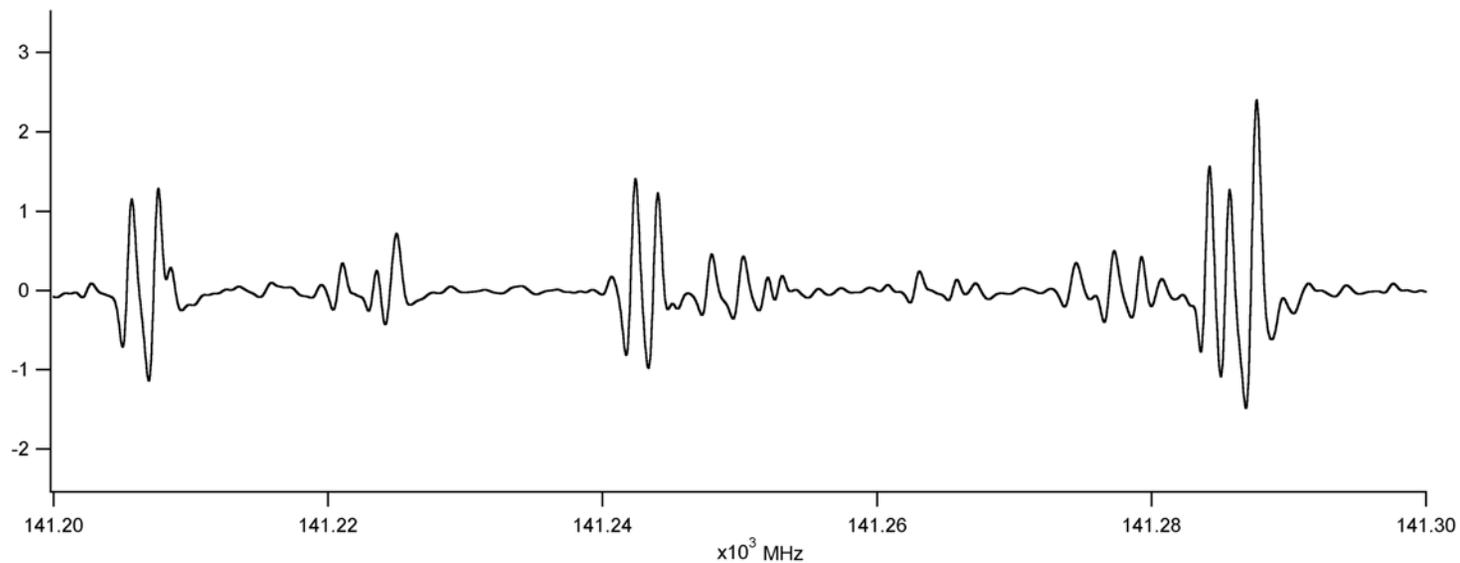


# EXPANDED FASST SPECTRUM OF $\text{ClONO}_2$

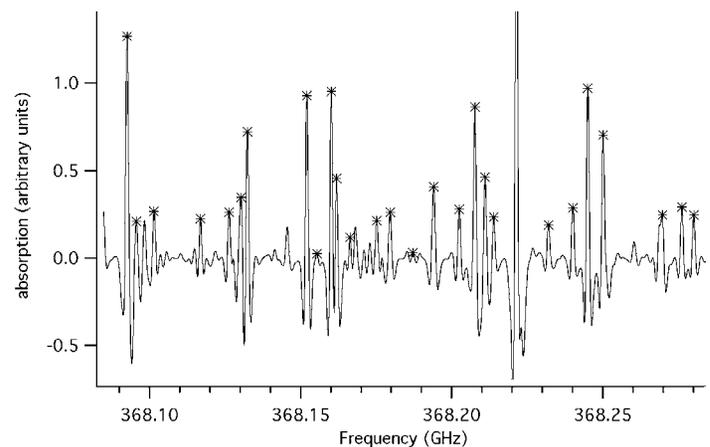
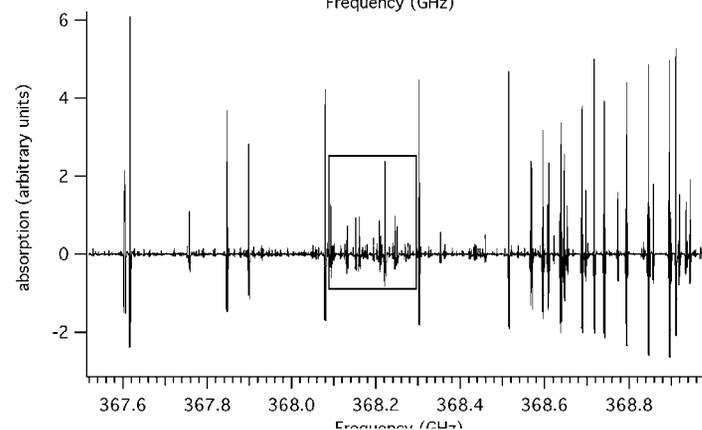
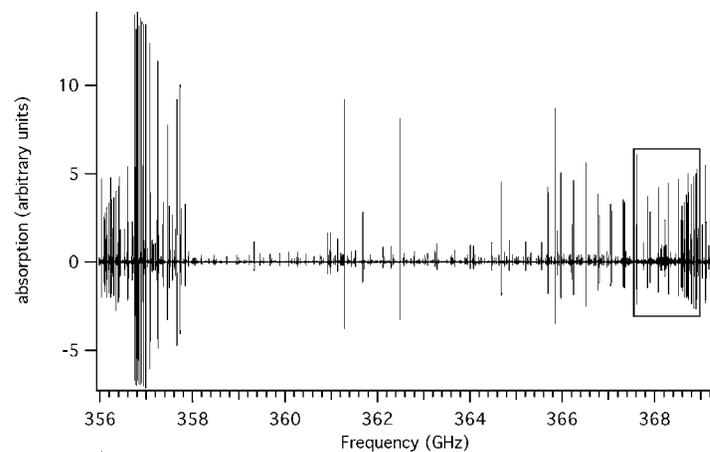


# FASSST SPECTRUM OF $\text{ClONO}_2$

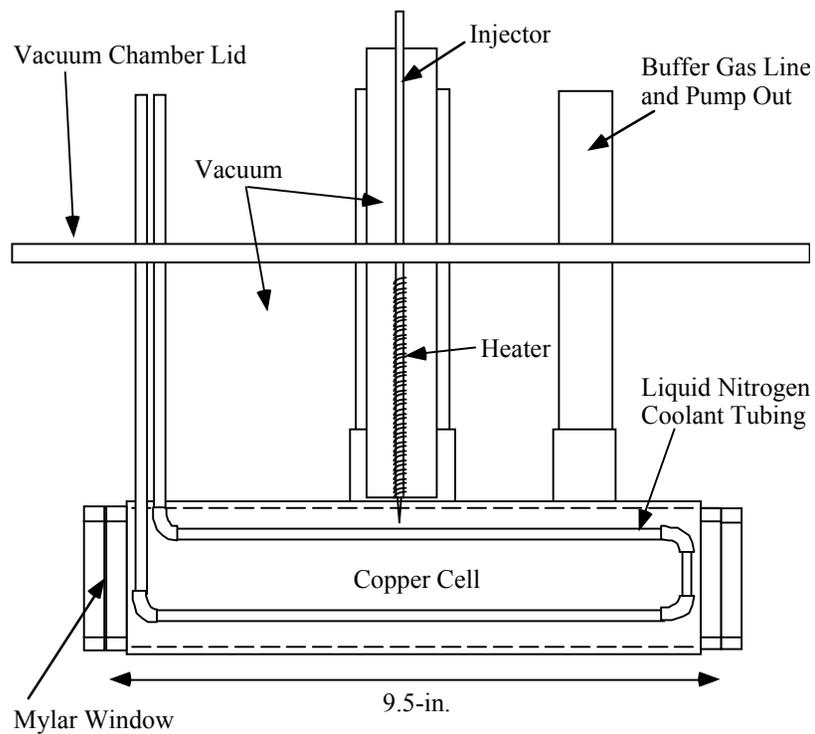
- 0.02 Second Measurement Time

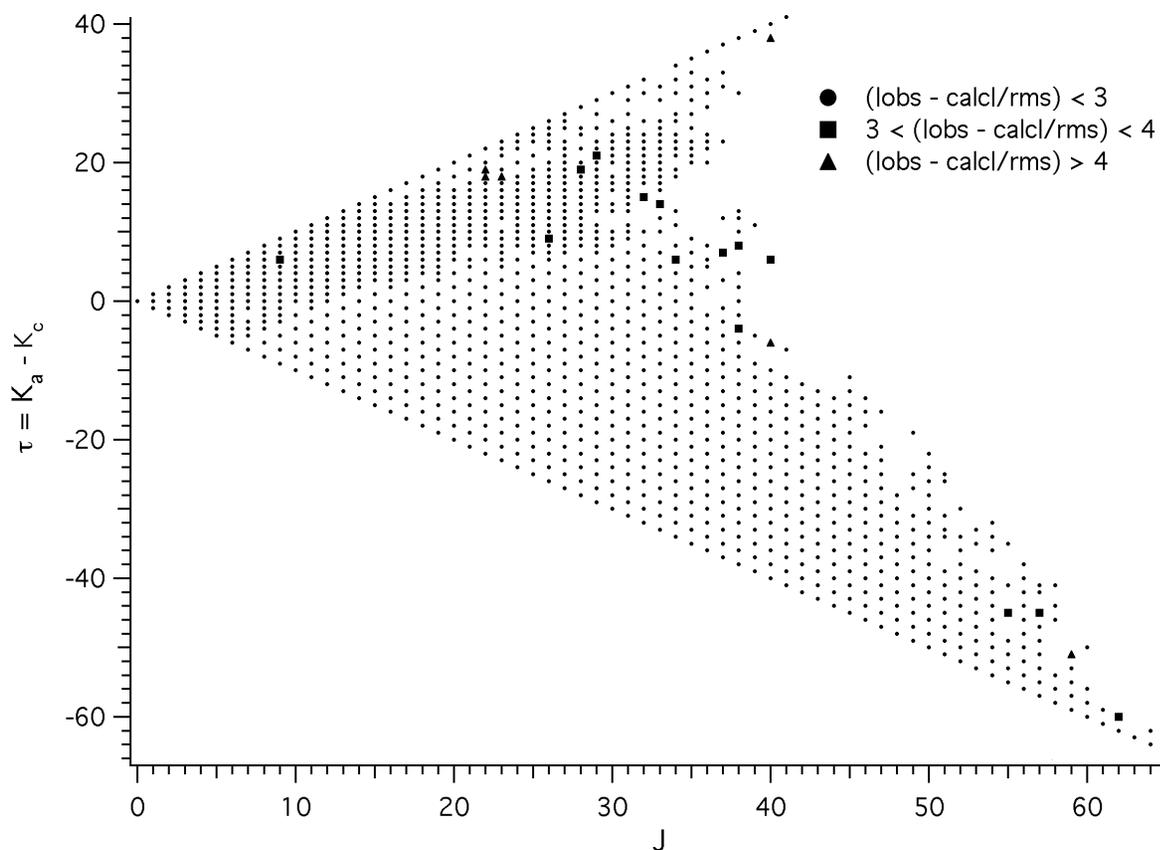


Series of expansions of a spectrum of  $\text{HNO}_3$  recorded by the FASSST spectrometer. The last expansion displayed shows 23 lines of the pure rotational spectrum of  $\nu_5$  and required 0.02 seconds to record. These lines are about 150 times weaker than lines in the ground vibrational state.

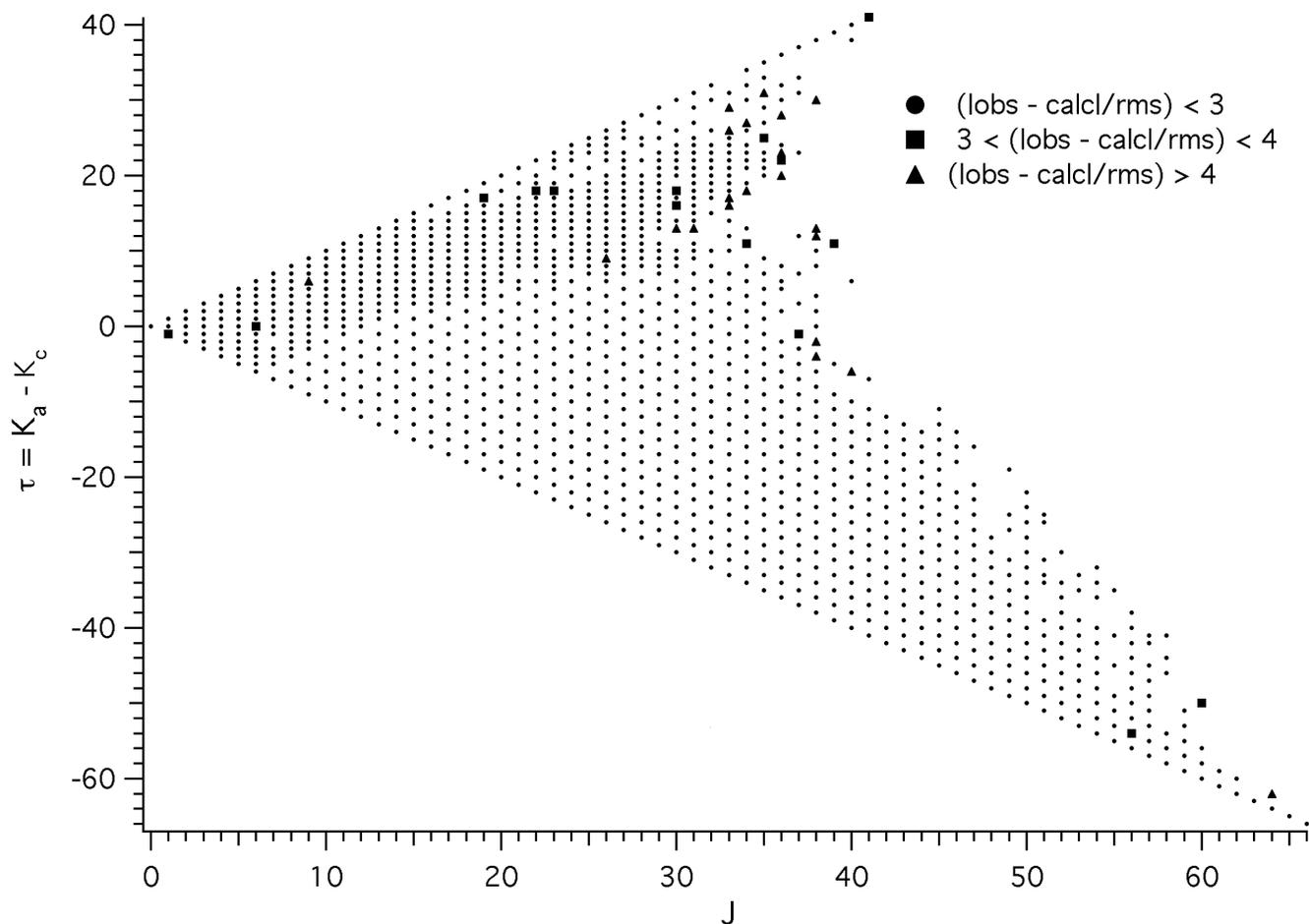


# SEPARATION OF ROTATIONAL AND VIBRATIONAL TEMPERATURES





Distribution of differences for a fit to the experimental infrared energy levels of  $\nu_5$  of  $\text{HNO}_3$ .



Distribution of differences between the experimental infrared energy levels of  $\nu_5$  and levels *predicted* from the microwave analysis.

## SPECIES OBSERVED BY IR SENSING BETWEEN 700 TO 1300 $\text{CM}^{-1}$

1.  $\text{H}_2\text{O}$ ,  $\text{CO}_2$ ,  $\text{O}_3$ ,  $\text{N}_2\text{O}$ ,  $\text{CH}_4$ ,  $\text{OCS}$ ,  $\text{NH}_3$ ,  $\text{C}_2\text{H}_6$
2.  $\text{HNO}_3$ ,  $\text{SF}_6$ ,  $\text{C}_2\text{H}_6$ ,  $\text{COF}_2$ ,  $\text{HCOOH}$ ,  $\text{CF}_4$
3.  $\text{CCl}_2\text{F}_2$  (CFC-12),  $\text{CHClF}_2$  (CFC-22),  $\text{CCl}_3\text{F}$  (CFC-11),  $\text{CCl}_4$
4.  $\text{ClONO}_2$ ,  $\text{HO}_2\text{NO}_2$ ,  $\text{N}_2\text{O}_5$

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Adapted from Brown, Farmer, Rinsland, and Zander

# THE ROTATION-VIBRATION DATA BASE

## WEIGHTED FITS OF MICROWAVE AND INFRARED DATA

Many infrared data points

Starting to be many microwave data points

Fits need to be at level of data, not derived levels

Fits need to include all relevant data

## LINewidths: EXPERIMENT AND THEORY

No 'Spectroscopy-like' Fitting Theory  
--> accuracy?

Gamache, Hartmann, Rosenmann Survey

JPL/OSU HNO<sub>3</sub> Intercomparison

## IR and MW LINEWIDTHS

Function of J?    light ( $\text{H}_2\text{O}$ ):                    yes  
                          medium ( $\text{HNO}_3$ ):    self - yes  
                          heavy (CFC's):                    ?

Function of  $\nu$ ?                    ?

A combined MW/IR broadening data base?

For which molecules?

For which states?

# CONTINUA

**Dimers and/or Polymers**

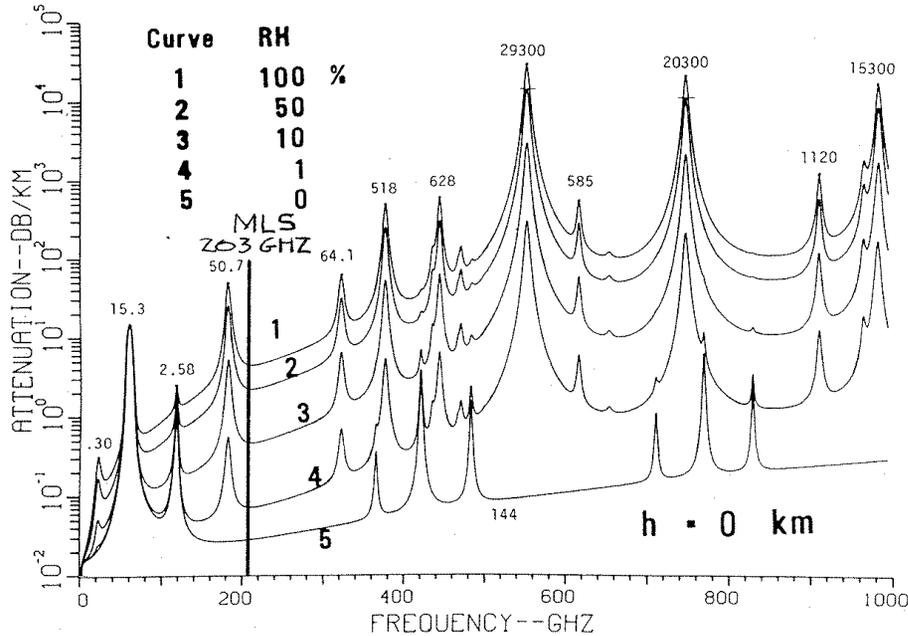
**Nearest Neighbor Lineshape Effects**

**Far Wing Lineshape Effects**

**Anomalous Absorption**

**Collision Induced Absorption**

# I. OVERVIEW



MLS Function  $\beta = Ap^2\left(\frac{300}{T}\right)^B$

Dry Air  $A = 7.30 \cdot 10^{-9} km^{-1} hPa^{-2}$   
 $B = 2.79$

Moist Air  $A = 5.67 \cdot 10^{-5} km^{-1} hPa^{-2}$   
 $B = 3.59$

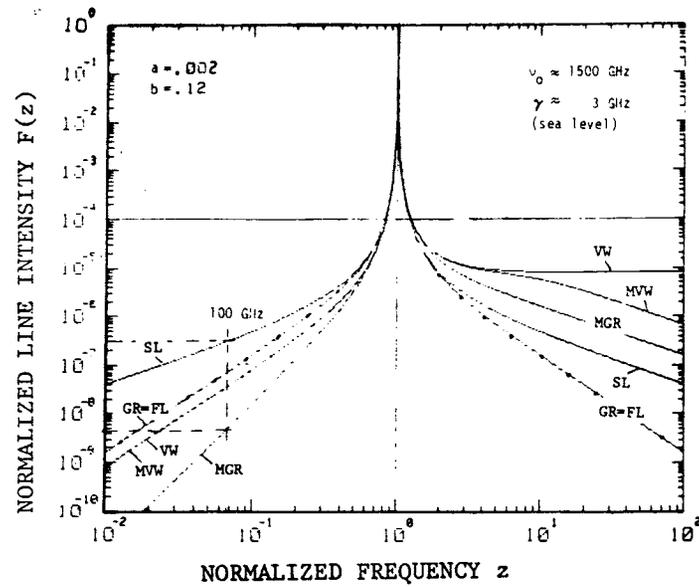
## Liebe Function

Dry Air  $\alpha(db/km) = 2.5 \cdot 10^{-11} f^2 p^2 \left(\frac{300}{T}\right)^{4.5}$

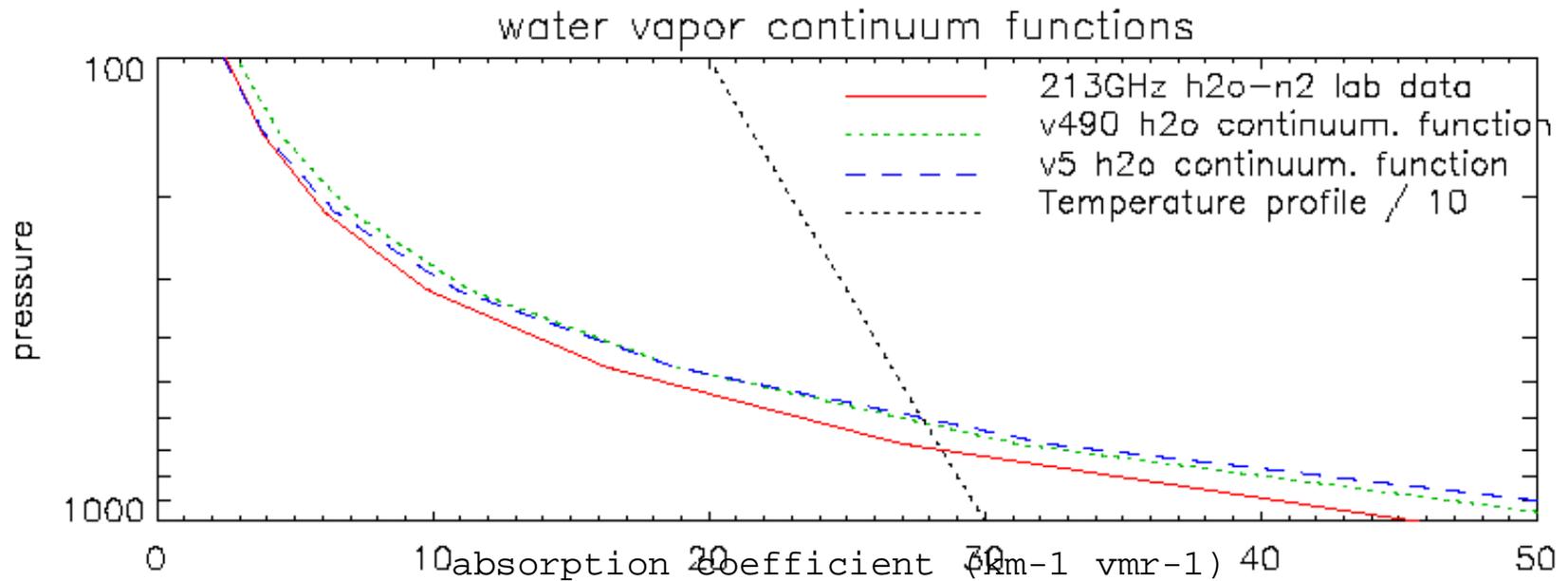
Moist Air  $\alpha(db/km) = 2.5 \cdot 10^{-7} p e f^2 \left(\frac{300}{T}\right)^{2.5} + 9.7 \cdot 10^{-6} e^2 f^2 \left(\frac{300}{T}\right)^{5.5}$

$p$  = partial pressure of dry air (kPa)  
 $e$  = partial pressure of water (kPa)  
 $f$  = frequency (GHz)

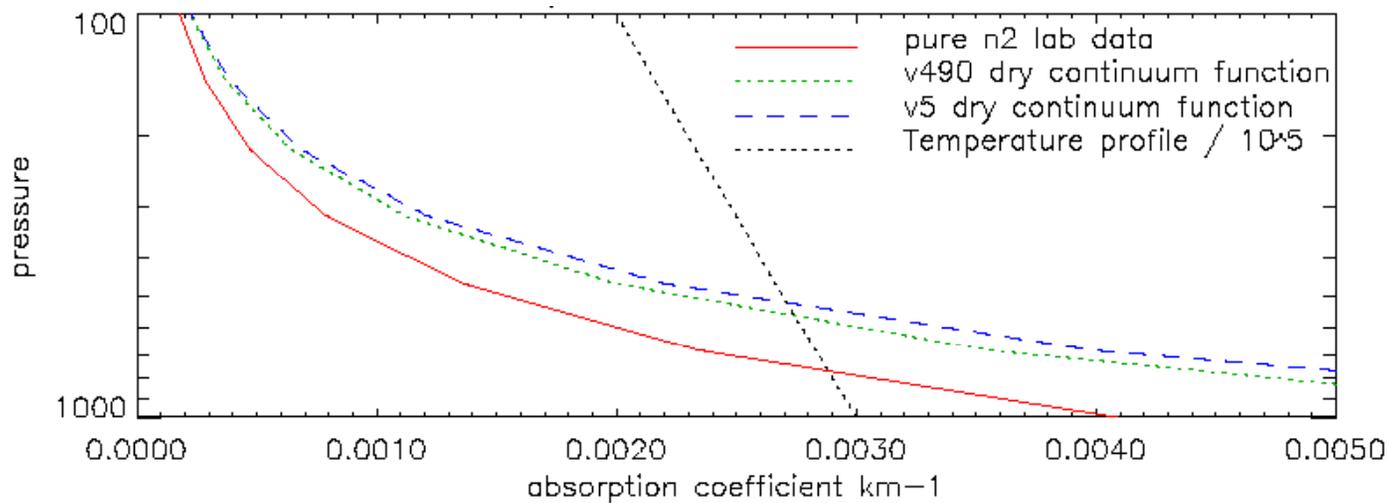
# Lineshapes



# Comparison of 'laboratory' and field measurements of the moist air continuum (from Read et al.)



## Comparison of 'laboratory' and field measurements of the dry air continuum (from Read et al.)



## II. THE PHYSICS

### Moist Air

Observed excess in atmospheric windows over simple

Lorentzian

Lineshape from nearby lines?

Far wings of very strong lines at higher frequencies?

Stable species (e. g. dimers)?

Metastable states, non-equilibrium physics?

More exotic physics?

### Dry Air

Collision induced due to  $O_2$  and/or  $N_2$



## III. THE EXPERIMENTAL ISSUES

### Moist Air Continuum:

What is in the cell?

Monomers, clusters, nonequilibrium species

(What is in the atmosphere?)

### Dry Air Continuum:

Small effect in the laboratory

What is in the cell?

Contamination from Moist Air Continuum

## IV. THE EXPERIMENT

10 m Fabry-Perot Cavity

$Q \sim 10^7$ ,  $\Delta\nu = 20$  kHz, EPL 1 - 10 km

Stainless Vacuum/Pressure Chamber

( $10^{-9}$  - 10 atm)



FASSST Scanning with skip-lock frequency stabilization

Continuous Scan:  $\sim 1000$  s, duty cycle  $\sim 10^{-3}$

Skip-lock Scan: adjustable scan time, duty cycle  $\sim 1$

## SUMMARY

**Due to advances in sensors and analysis**

**Significant Advances in IR and MW Field Data**

**Significant Advances in IR and MW Laboratory Data**

**We should pay careful attention to the**

**Integration of MW/IR data base**

**Transition between lineshape and continua**